

Nuclear physics, radioactive decay

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Electromagnetic Spectrum

The electromagnetic spectrum stretches from radio waves to gamma rays.

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Organs	Tissues	Cells	Proteins	mRNA	DNA
Computer Tomography (CT) Ultrasound (US) Magnetic Resonance Imaging (MRI)		Optical Imaging Nuclear Imaging (PET/SPECT)			

large availability of highly specific tracers

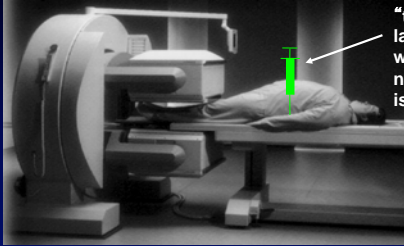
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highly sensitive measurement of a large variety of biological processes

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What is Radionuclide Imaging? (a.k.a. "Nuclear Medicine")

Standard Nuclear Emission Camera



Inject "tracer" labeled with a nuclear isotope

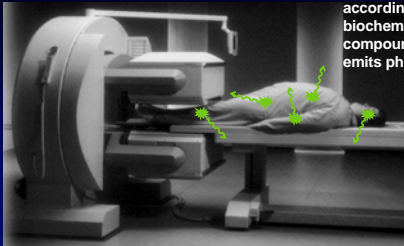


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What is Radionuclide Imaging?

Standard Nuclear Emission Camera



Tracer distributes according to biochemistry of compound and emits photons



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The Atomic Nucleus

•Z protons

•N neutrons

} Nucleons



•A = Z+N = mass number

•Nuclear composition of atomic element X: ${}^A X_Z$ or ${}_Z X^A$

•Same Z, different A \rightarrow Isotope (chemically identical)



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Forces Within the Nucleus

- *Strong* nuclear forces bind nucleons together
- *Weak* nuclear forces control certain nuclear decays
- *Electromagnetic* forces cause repulsion between protons
- *Binding energy* is the energy required to break apart nucleus into its individual constituents:

$$BE(A,Z)=[ZM_H+NM_N-M(A,Z)]c^2$$



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There are Tremendous Forces Within the Nucleus

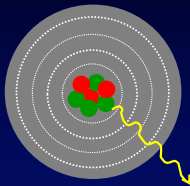


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What is a radionuclide?

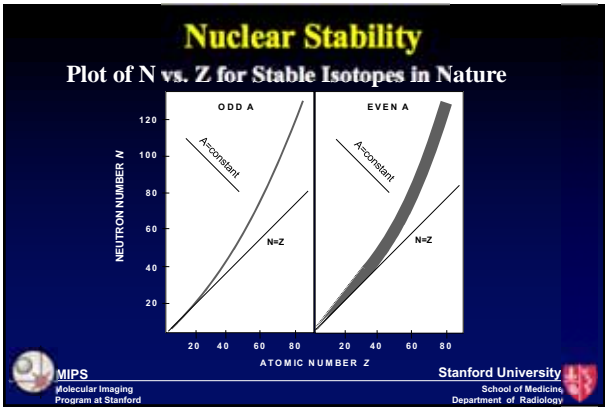
A nucleus in an energetic (“unstable”) condition that relaxes (“decays”) by emitting radiation in the form of energetic particles



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Relevant Modes of Nuclear Decay for Imaging

•**Gamma Ray Emission (Isomeric Transition):**
 Nucleons (protons and neutrons) rearrange themselves to a less energetic, more stable configuration, giving off electromagnetic energy $E_f - E_i - E_\gamma$ in the form of gamma (γ) rays. For imaging, need *metastable* parents. Z does not change.

$${}^A X_Z \rightarrow {}^A X_Z + \gamma \quad (\text{or internal conversion electrons}) \quad (\neq \text{x-ray or Auger } e^-)$$

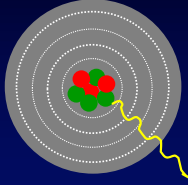
Example nuclei that result in γ emission: ${}^{111}\text{In}$, ${}^{123}\text{I}$, ${}^{125}\text{I}$, ${}^{99m}\text{Tc}$, ${}^{201}\text{Tl}$

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Gamma ray decay (isomeric transition)

A nucleus in an energetic ("unstable") condition that relaxes ("decays") by emitting gamma rays



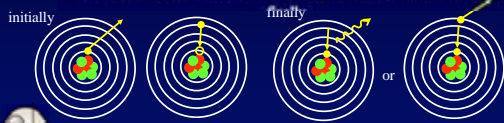
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Internal conversion competes with gamma ray decay

Excited nucleus relaxes by ejection an inner shell electron, which leaves an inner shell electron vacancy which is filled by outer shell electrons and the excess energy is either emitted as:

1. x-ray (Characteristic radiation) or
2. Transferred to an outer shell electron (Auger electron)



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Relevant Modes of Nuclear Decay for Imaging

•Beta Decay (Isobaric Transition):

Nucleons transform into each other to obtain a more stable configuration. This involves the emission or absorption of a beta (β) particle.

Examples: $n \rightarrow p + e^- + \bar{\nu}$ (electron emission); $p \rightarrow n + e^+ + \nu$ (positron emission); $e^- + p \rightarrow n + \nu$ (electron capture)

For Beta Emission, $E_\beta + E_\nu = M_N - M_P$. Z changes.



For electron capture: $e^- + {}^A X_Z \rightarrow {}^A Y_{Z-1} + \nu$ (+ x-ray or Auger e^-)

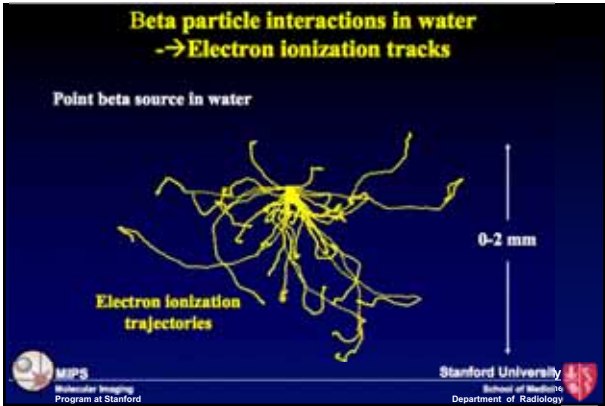
Example β -decaying nuclei: ${}^3\text{H}$, ${}^{14}\text{C}$, ${}^{32}\text{P}$, ${}^{11}\text{C}$, ${}^{13}\text{N}$, ${}^{15}\text{O}$, ${}^{18}\text{F}$

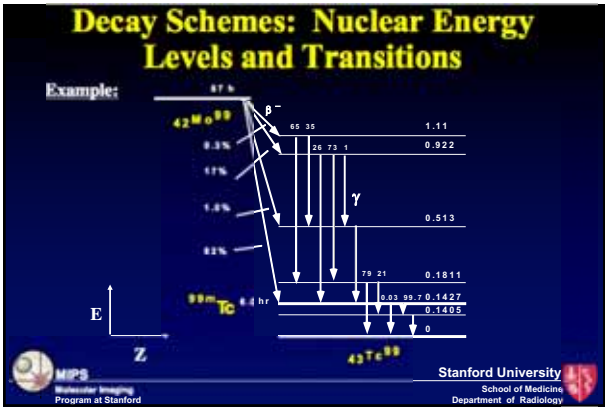
Example electron capture decaying nuclei: ${}^{111}\text{In}$, ${}^{201}\text{Tl}$

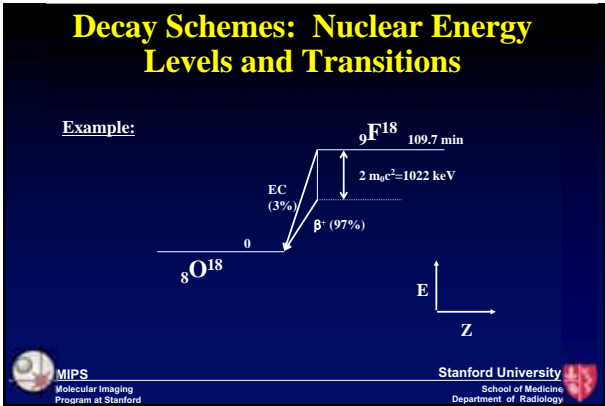


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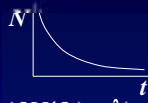
Nuclear Decay Probability

Radioactivity is a random process. If N_0 identical unstable parent nuclei were present initially, the most probable number N surviving after a time t is given by:

$$N = N_0 e^{-\lambda t}$$

λ = decay constant

$$\text{Activity} \equiv -dN/dt = \lambda N_0 e^{-\lambda t} = \lambda N = (dN/dt)_0 e^{-\lambda t}$$



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Marie Curie



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Radioactivity

Exponential radioactive decay

If $(dN/dt)_0$ is the initial activity of a radioactive source, then the activity after a time t will be given by:

$$dN/dt = (dN/dt)_0 e^{-\lambda t}$$

(Ci or Bq)
1 Ci = 3.7×10^{10} dis/sec
= 3.7×10^{10} Bq

λ = decay constant = $\ln 2 / \tau_{1/2}$

$\tau_{1/2}$ = half-life of radionuclide



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Some Common Radionuclides Used in "Nuclear Medicine"

Isotope	Energy (keV)	Half-life	Production
⁶⁷ Ga	93, 185, 300	78 hours	Cyclotron
¹¹¹ In	172, 247	67 hours	Cyclotron
^{113m} In	392	100 minutes	Cyclotron
¹²³ I	159	13.2 hours	Cyclotron
¹³¹ I	364	8 days	Fission
^{99m} Tc	140	6 hours	Generator
²⁰¹ Tl	80, 167	73 hours	Cyclotron
¹³³ Xe	80	5 days	Fission
¹⁸ F	511 keV	110 min.	Cyclotron
¹¹ C	511 keV	20.4 min.	Cyclotron

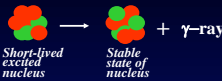


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Nuclear decays of interest for imaging lead to generation of high energy photons

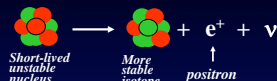
Gamma decay: Nuclear de-excitation



Example: $^{99m}\text{Tc} \rightarrow ^{99}\text{Tc} + \gamma$



Positron decay: Nuclear transmutation



Example: $^{18}\text{F} \rightarrow ^{18}\text{O} + e^+ + \nu$



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